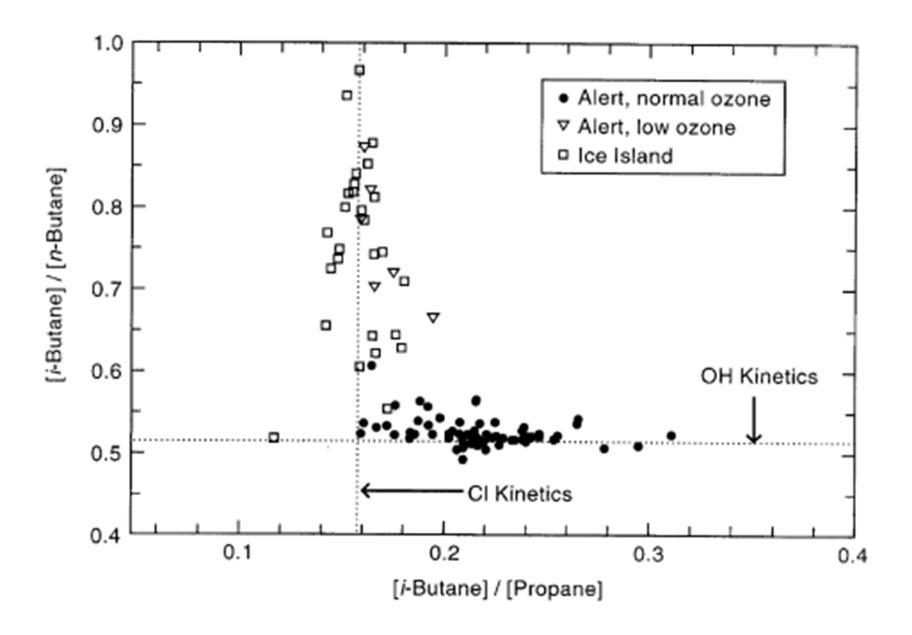
"Perspective on Halogens"

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What things do we know?



We know chlorine atom chemistry is occurring and it is important! (Jobson et al., JGR, 1994)

OASIS09 –

Measurements of "everything that matters"

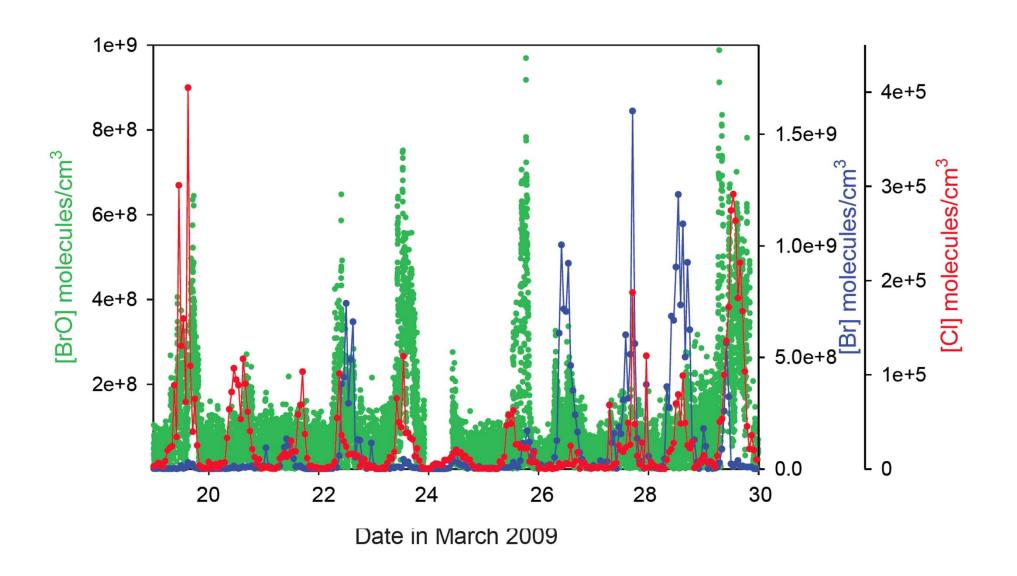
 $[Cl]_{ss} = (2J[Cl_2] + J[BrCl] + J[ClO] + k[Cl_2][OH] + k[ClO][OH] + k[ClO][NO] + k[ClO][ClO])$

 $(k[HO_2] + k[O_3] + k[MEK] + k[CH_4] + k[C_2H_2] + k[C_2H_4] + k[C_2H_6] + k[C_3H_6] + k[C_3H_8] + k[iC_4H_{10}] + k[nC_4H_{10}] + k[HCHO] + k[CH_3CHO])$

 $[Br]_{ss} = (2J[Br_2] + J[BrCl] + J[BrO] + J[HOBr] + k[Br_2][OH] + k[BrO][NO] + k[BrO][ClO] + 2k[BrO][BrO] + k[BrO][OH])$

 $(k[HO_2] + k[O_3] + k[NO_2] + k[HCHO] + k[CH_3CHO] + k[C_2H_2] + k[C_2H_4] + k[C_3H_6])$

OASIS09 Measured and Calculated Radical Concentrations



For "typical" daytime peaks of [CI] = $1x10^5$ and [Br] = $4x10^8$, [BrO] = $6x10^8$

Calculated rate of O_3 destruction by Br = 0.26ppb/hr.

Calculated rate of O_3 destruction by CI = 0.0.07ppb/hr.

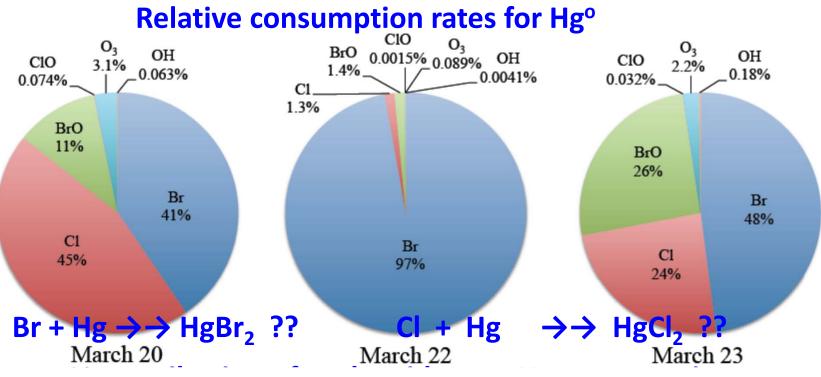
21%

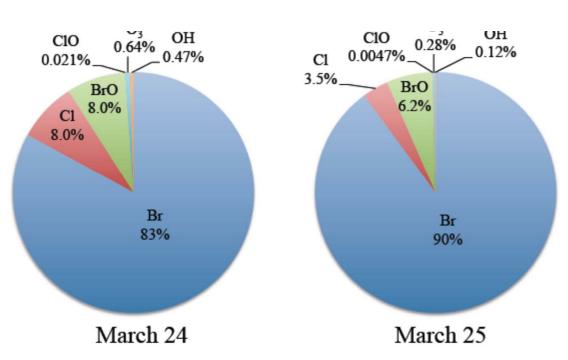
1x10⁵ cm⁻³ Cl represents a large oxidizing power.

For example, at T=245, for 1x10⁵ Cl and 1x10⁶OH

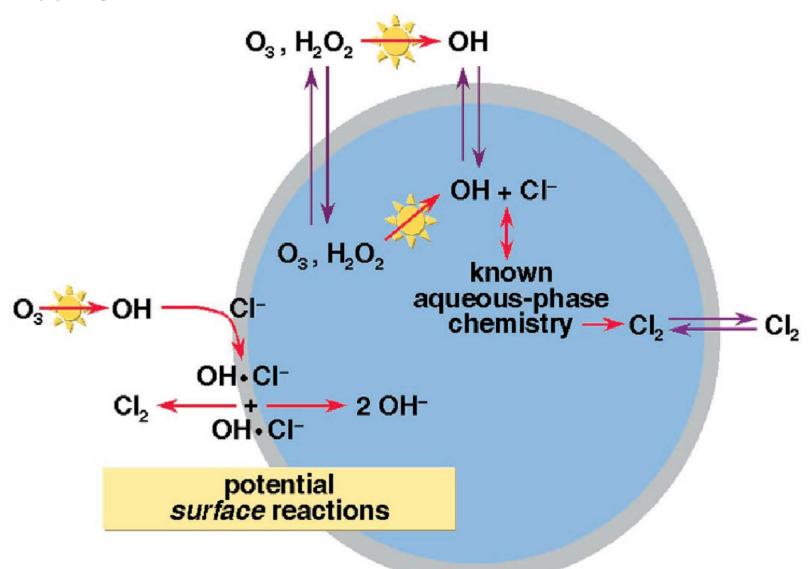
The first order k for consumption of n-butane is 1.74x10⁻⁶s⁻¹ for OH (6.6 days) and

 $2.0x10^{-5}s^{-1}$ for Cl (0.56 days)

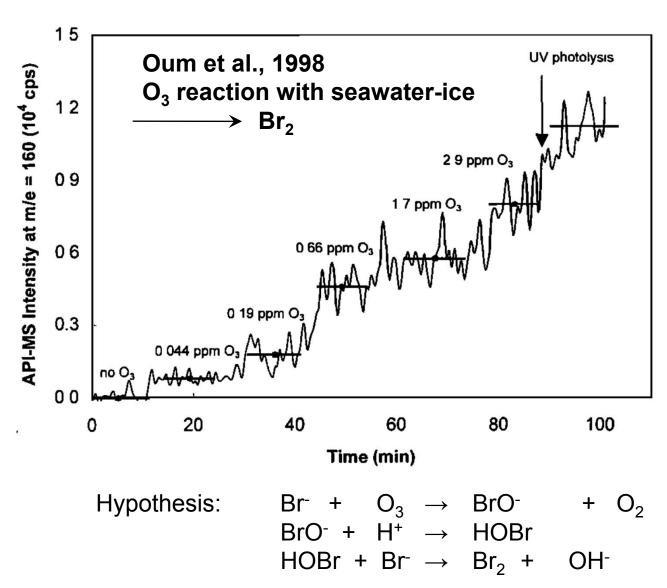


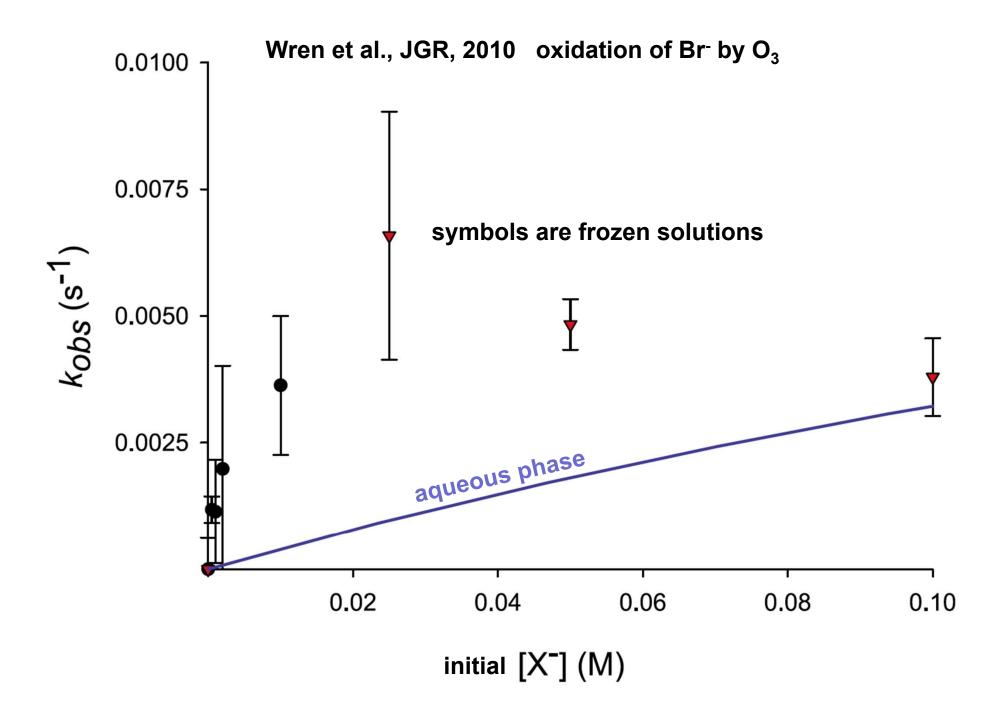


We know that you can make Cl_2 by irradiating sea salt aerosols with 254nm radiation in the presence of O_3 . Knipping et al., 2000

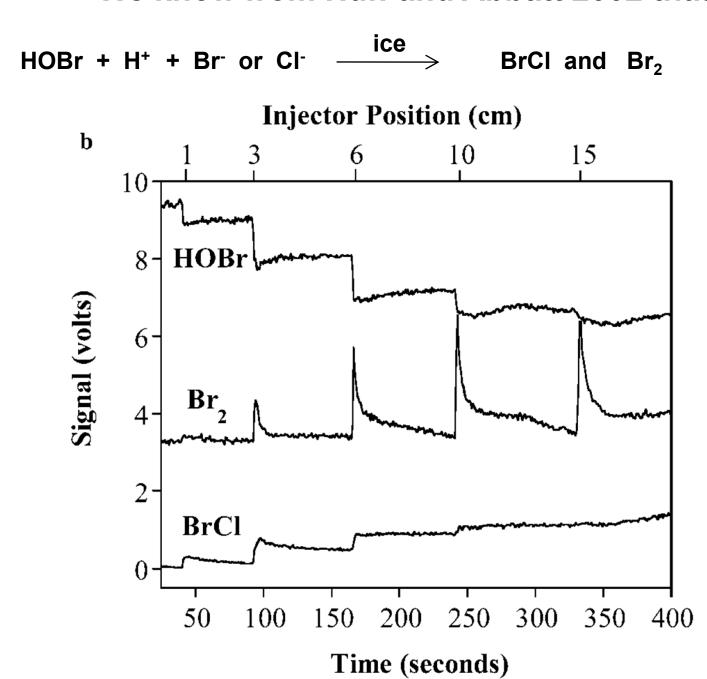


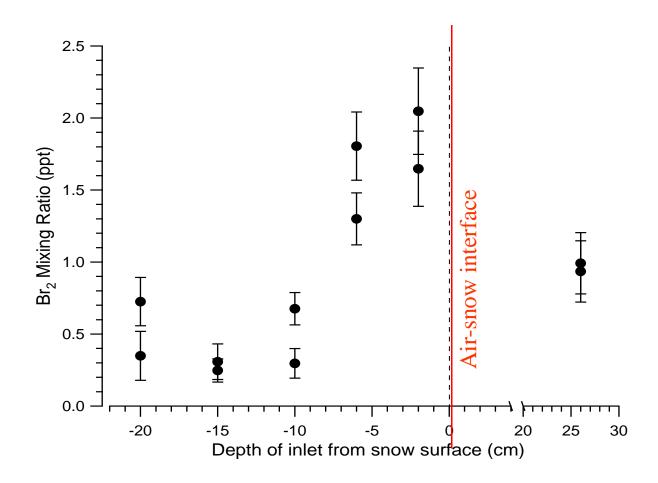
We know, e.g. from Oum et al. and Wren et al. that Br⁻ at surfaces can be oxidized by O₃ to produce Br₂ (snow/sea ice/sea salt aerosols)





We know from Huff and Abbatt 2002 that





We know that Br₂ is produced in saline snowpacks in the dark.

Foster et al., Science, 291, 471-474, 2001.

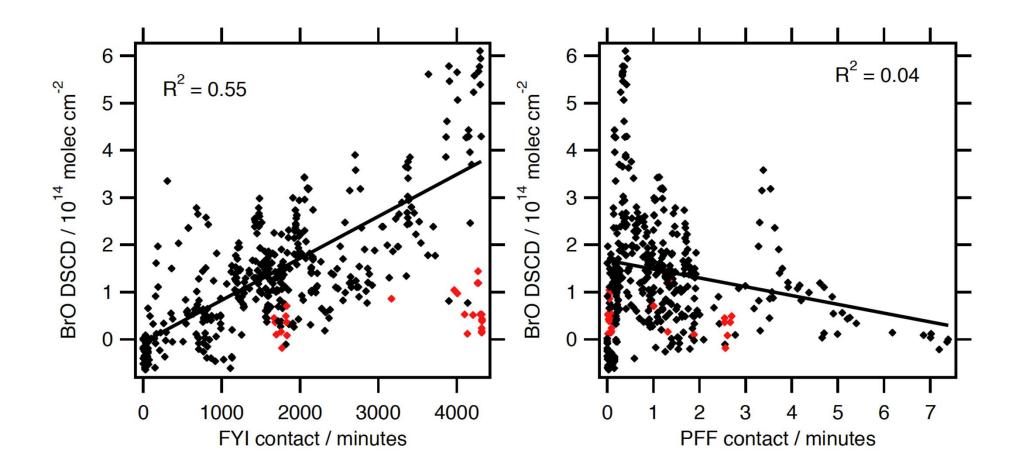
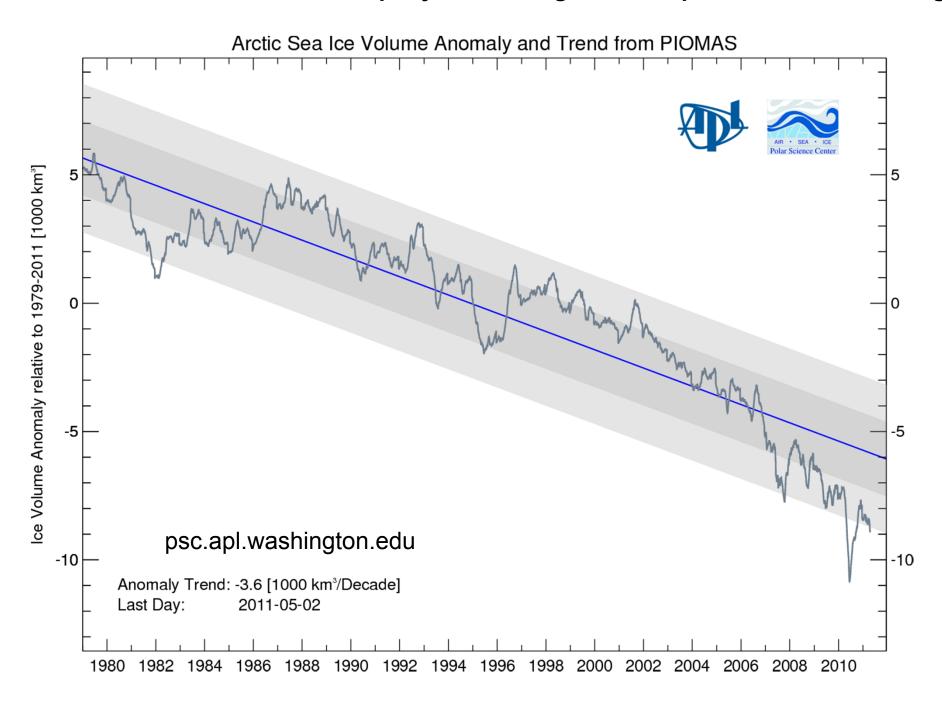
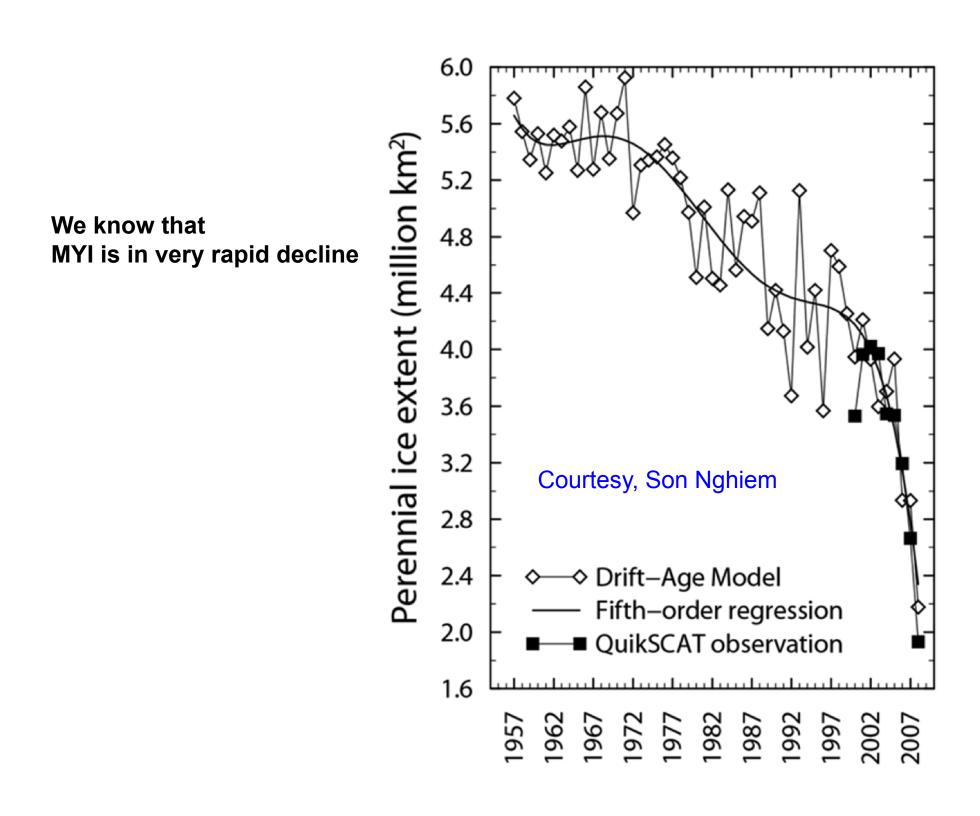


Fig. 3. Correlation plots of BrO versus first-year sea-ice contact (left) and potential frost flowers contact (right). Data colored in red occurred when ozone <1 ppbv and were ignored from the correlation analysis.

We know that BrO seems to correlated with First Year Sea Ice Simpson et al., 2006

We know that sea ice mass is rapidly decreasing, and that process is accelerating.



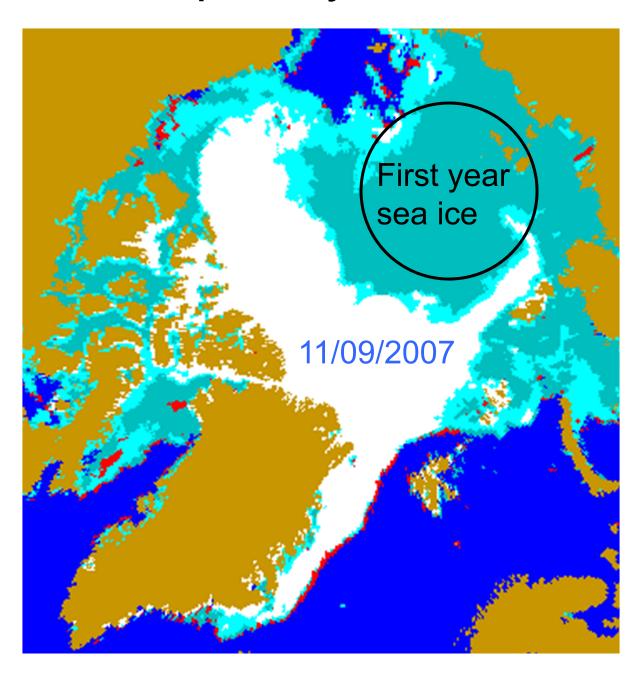


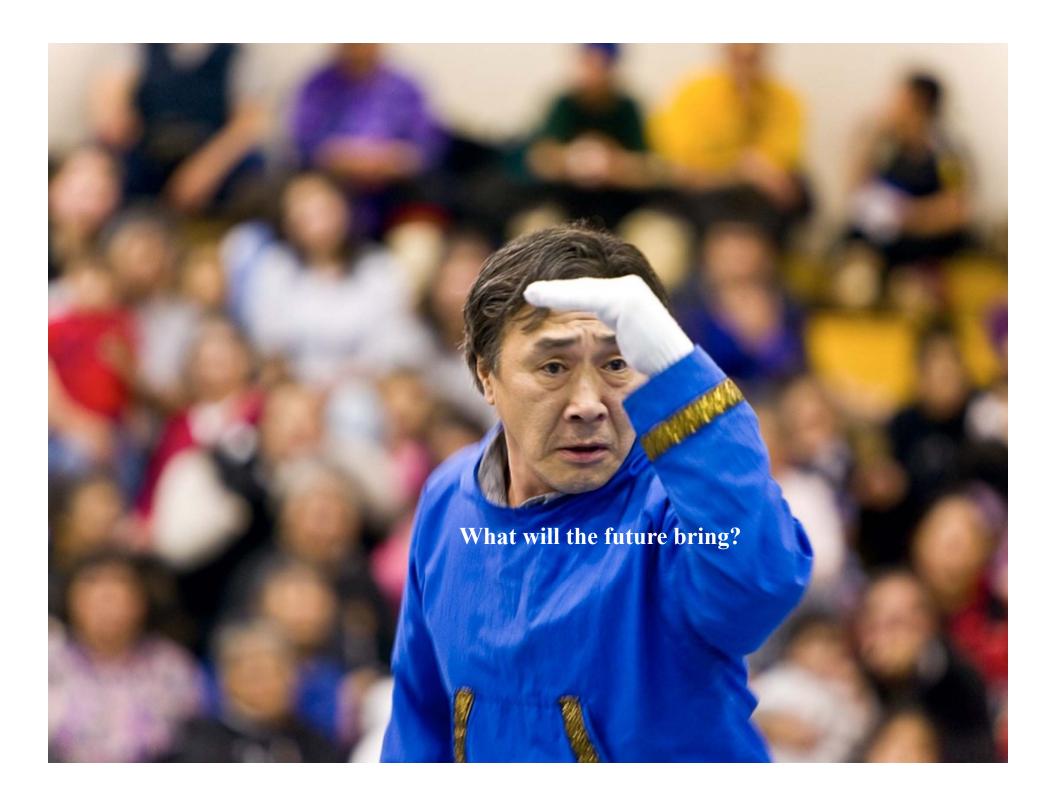


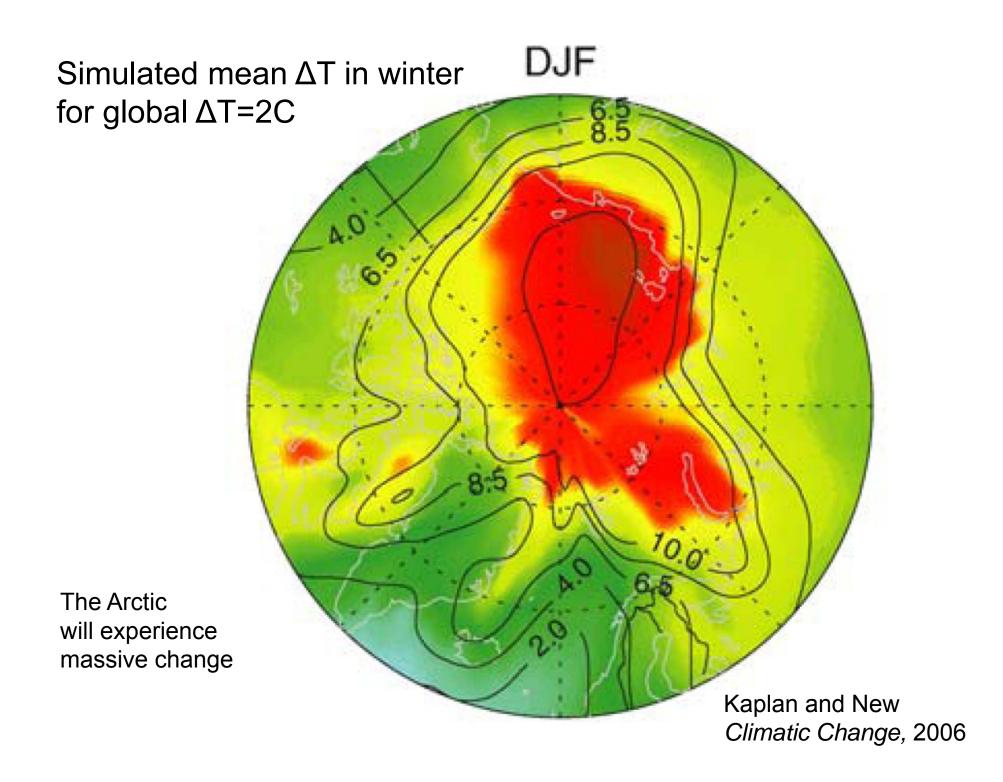
It will be replaced by First Year Ice



Courtesy Son Nghiem

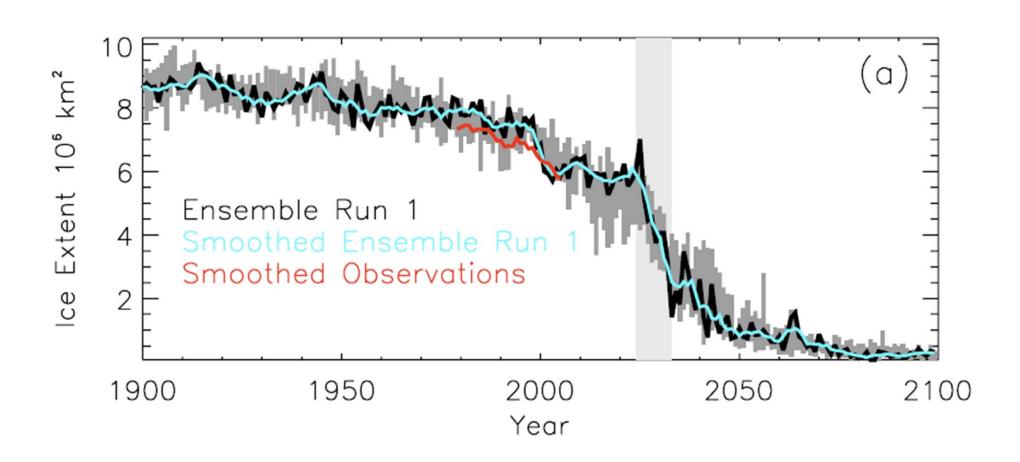




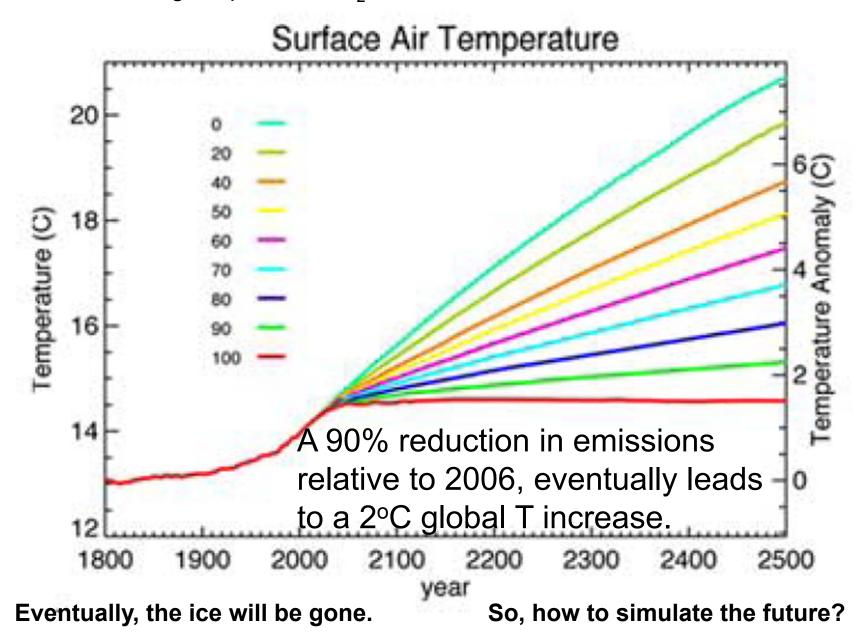


Holland et al., GRL, 2006

Arctic Sea Ice Extent Simulations



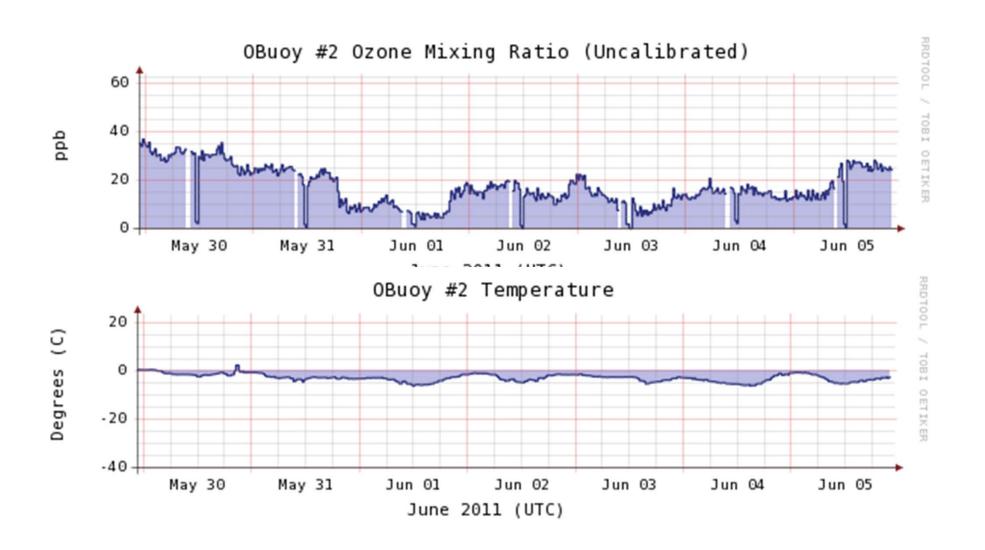
Weaver et al. (UVic) GRL 2007 Climate Change impacts of CO₂ emissions reductions relative to 2006 levels



How will an ice-free Arctic impact atmospheric composition?

- •More O₃ at the surface?
- Smaller oxidation capacity
- •Slower rate of production of CCN?
- •But more water vapor? So, more cloud cover?
- •Slower rate of Hg oxidation product inputs?
- •Can we investigate these questions from measurements over polynyas?
- •How to simulate this world?

We do know that ODEs can occur at relatively high T. See, e.g. OBuoy number 2, currently in the Beaufort Sea:



We think we know that this chemistry occurs on surfaces. For ice, for lack of a good term, let's call it the QLL.

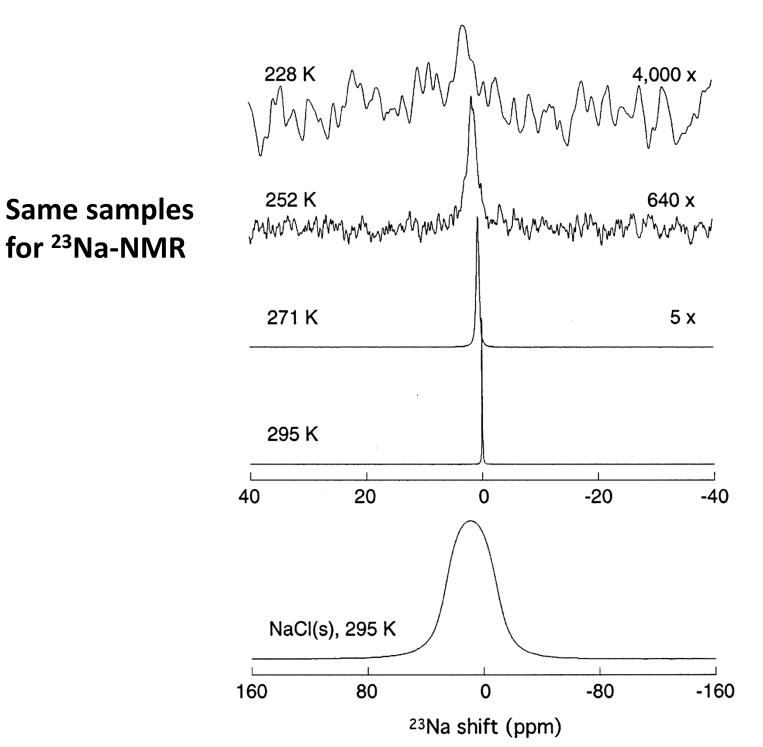
We think we know that this surface is a god-awful horrible mess.

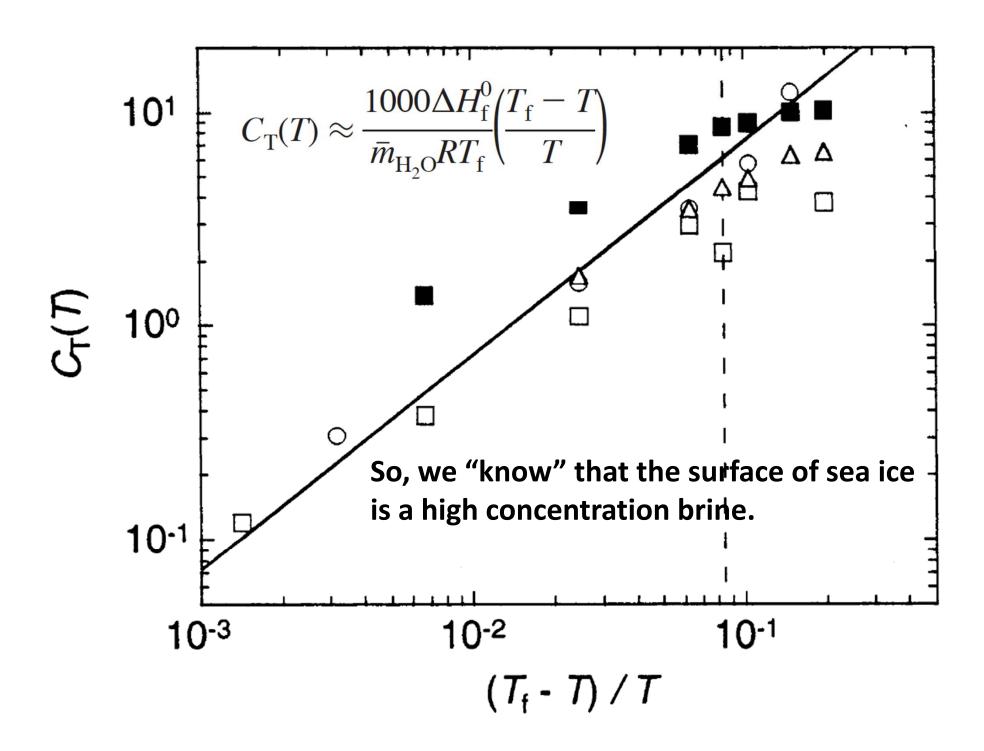
228 K 400,000 x 252 K 50,000 x 271 K 125 x 295 K 8 10 6 4 2 0

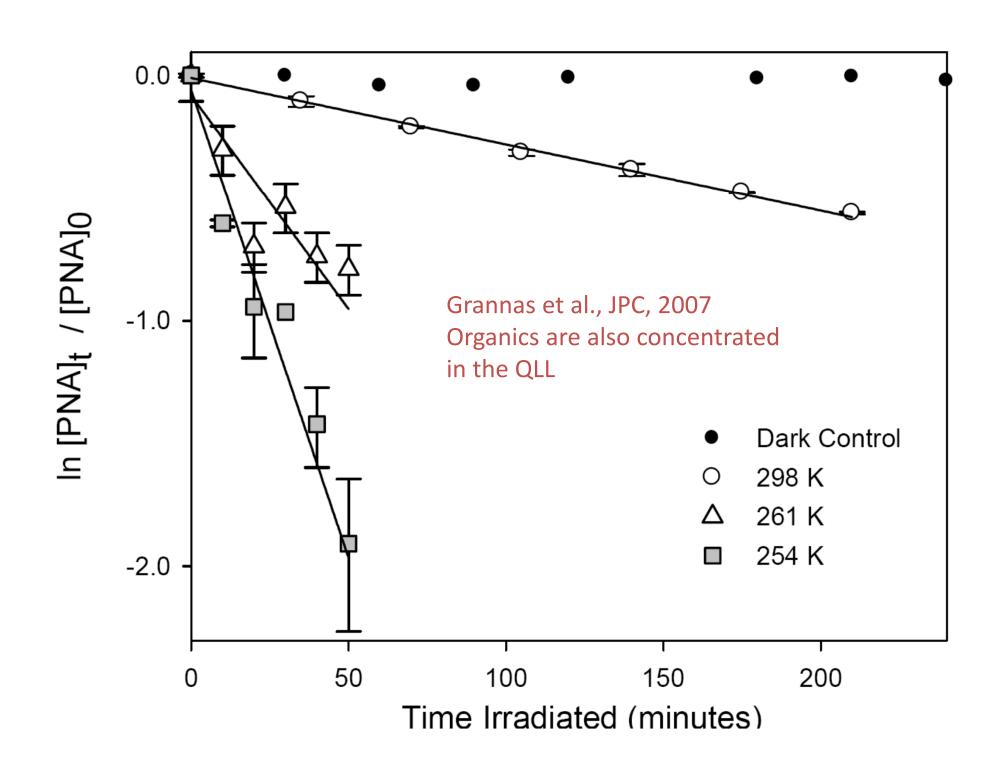
Figure 2. ¹H NMR spectra of a brine solution ([NaCl] = 0.500 M) at different temperatures. Relative intensity scales are given by the factors on the right. Each spectrum is an average of four scans. The time domain signals were apodized, but in no case was the width of the apodization function more than 20% of the intrinsic width of the ¹H line.

¹H shift (ppm)

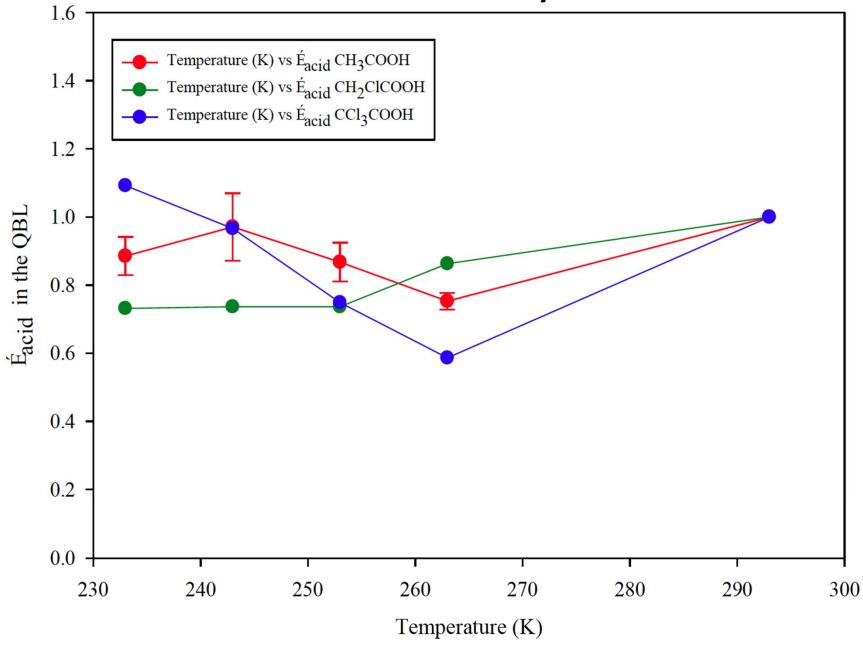
Cho et al., *J. Phys. Chem.*, 2002.







¹³C NMR results for carboxylic acids

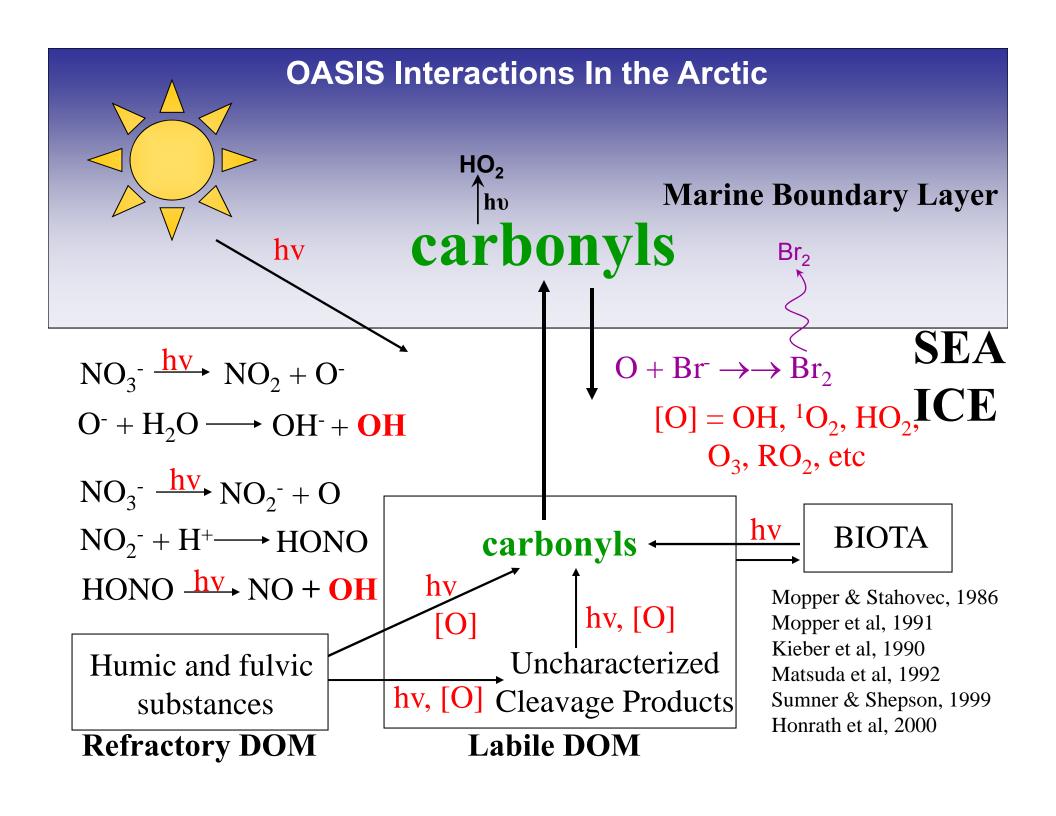


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Table 1. Most Abundant Molecular Formulas Identified in Both Ice Core Samples^a

1950 Core Sample						1300 Core Sample		
Formula	Error in Formula Mass, ppm	Peak Abundance	Formula	Error in Formula Mass, ppm	Peak Abundance	Formula	Error in Formula Mass, ppm	Peak Abundance
$C_{28}H_{40}O_4$ $C_{29}H_{42}O_4$	0.3 0.0	100 47.9	C ₂₅ H ₂₀ SO C ₂₅ H ₂₂ SO	$-0.1 \\ -0.2$	0.06 0.04	C ₂₈ H ₄₀ O ₄ C ₂₉ H ₄₂ O ₄	$0.2 \\ -0.2$	100 52.4
$C_{26}H_{36}O_4$	0.2	45.4	$C_{25}H_{24}SO$	0.1	0.06	$C_{26}H_{36}O_4$	0.2	35.7
$C_{24}H_{18}O_{10}$	0.4	26.5	$C_{25}H_{26}SO$	0.0	0.04	$C_{19}H_{34}O_{6}$	-0.3	14.3
$C_{24}H_{44}O_{12}$	0.2	24.2	$C_{25}H_{20}SO_2$	-0.5	0.17	$C_{26}H_{48}O_{13}$	0.2	13.7
$C_{26}H_{48}O_{13}$	0.3	21	$C_{25}H_{22}SO_2$	0.6	0.11	$C_{24}H_{44}O_{12}$	0.2	12.7
$C_{27}H_{26}O_{13}$	0.4	10.4	$C_{25}H_{24}SO_2$	0.4	0.07	$C_{22}H_{40}O_{11}$	0.1	10.7
$C_{30}H_{44}O_4$	0.4	8.3	$C_{25}H_{20}SO_3$	0.2	0.09	$C_{29}H_{46}O_{7}$	0.1	9.9
$C_{24}H_{40}O_4$	0.1	7.8	$C_{25}H_{22}SO_3$	-0.6	0.06	$C_{30}H_{44}O_4$	0.1	8.7
$C_{29}H_{46}O_{7}$	0.2	5	$C_{25}H_{24}SO_3$	0.5	0.06	$C_{20}H_{36}O_{10}$	0.2	7.7
$C_{26}H_{42}N_4O_4$	0.2	4	$C_{26}H_{24}SO_3$	0.5	0.10	$C_{27}H_{42}O_6$	-0.2	6.9
$C_{29}H_{50}O_4$	0.0	3.6	$C_{26}H_{26}SO_3$	0.8	0.06	$C_{35}H_{58}O_{10}$	0.6	6.0
$C_{25}H_{30}O_{11}$	0.1	3.2	$C_{27}H_{22}SO_3$	0.0	0.13	$C_{24}H_{40}O_4$	-0.1	5.9

The QLL for all environmental samples will be a god-awful mess!



We do know from a number of people, like Domine et al., and Ammann et al., and Kahan et al., that the QLL water does not behave like aqueous water! So, to simulate this world, we have to address the things we don't know.

Some of the things we don't know

•What the actual chemistry is that produces Br₂ and Cl₂

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•For the chemistry, e.g. HOBr + H^+ + Br^- \rightarrow Br_2 + H_2O What rate constants do we use?
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•For the important equilibria, like:

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Br_2 (QLL) \leftrightarrow Br_2 (gas) What equilibrium constants do we use?
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- •Where are the reactants? At the ice/QLL interface? At the QLL/air interface?
- •What are the diffusion coefficients?
- •What are the organic reactants and how much is there?
- •Is microbiology important?

So, in other words:

We don't understand our container.

We don't understand the bulk solvent.

We don't know what our reactants are.

We don't know where they are in the container or what their concentrations are.

Simulating the current polar surface photochemistry is very difficult. Simulating it for the future will be even harder. Given this....

I think we need to first have a really good answer to the following questions:

Why should we?

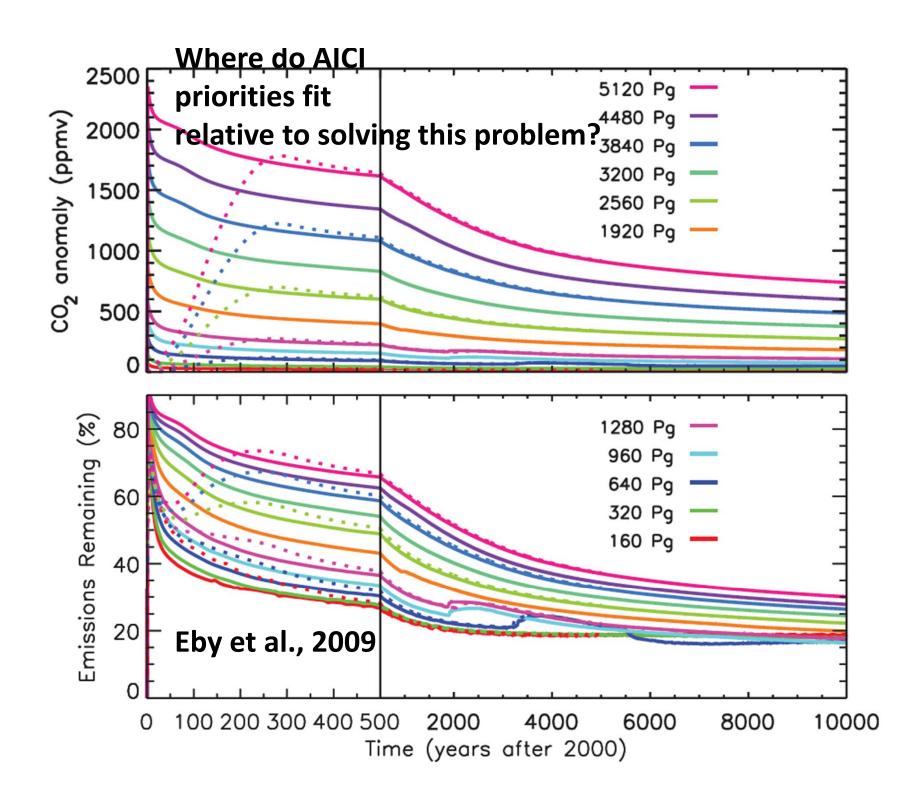
Why should we invest the effort?

Why is it important to be able to simulate O_3 at the poles in the future?

Or does the justification hinge on Hg?

Or something else?

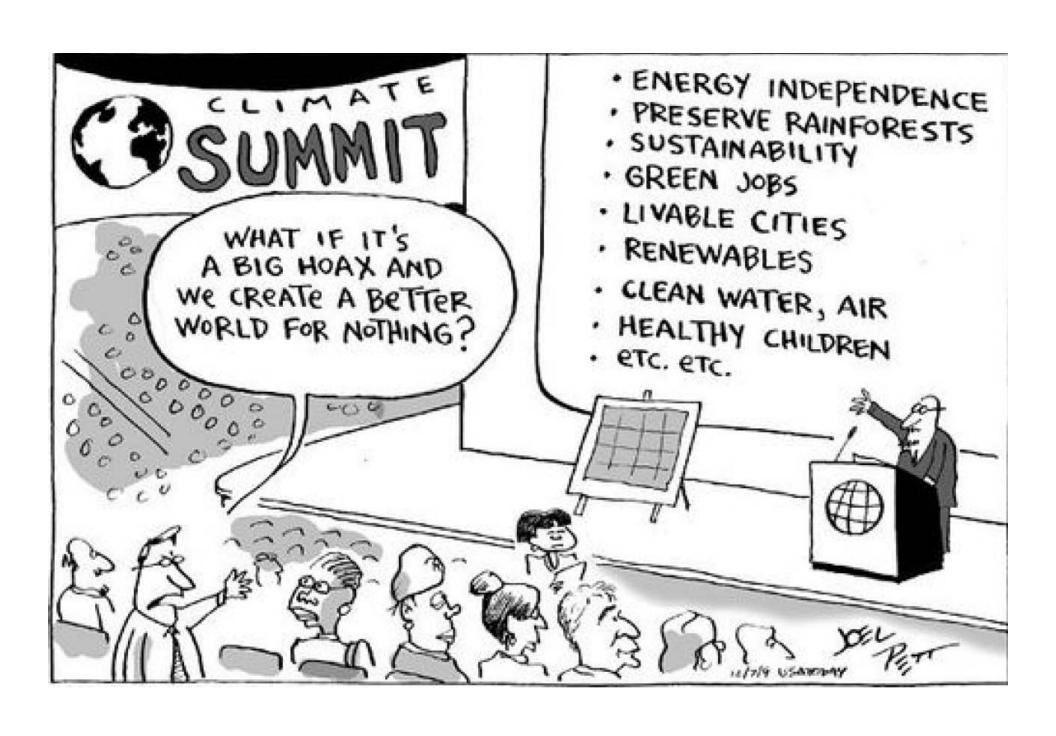
The answers to these questions influence what you work on, and how much funding the community gets.

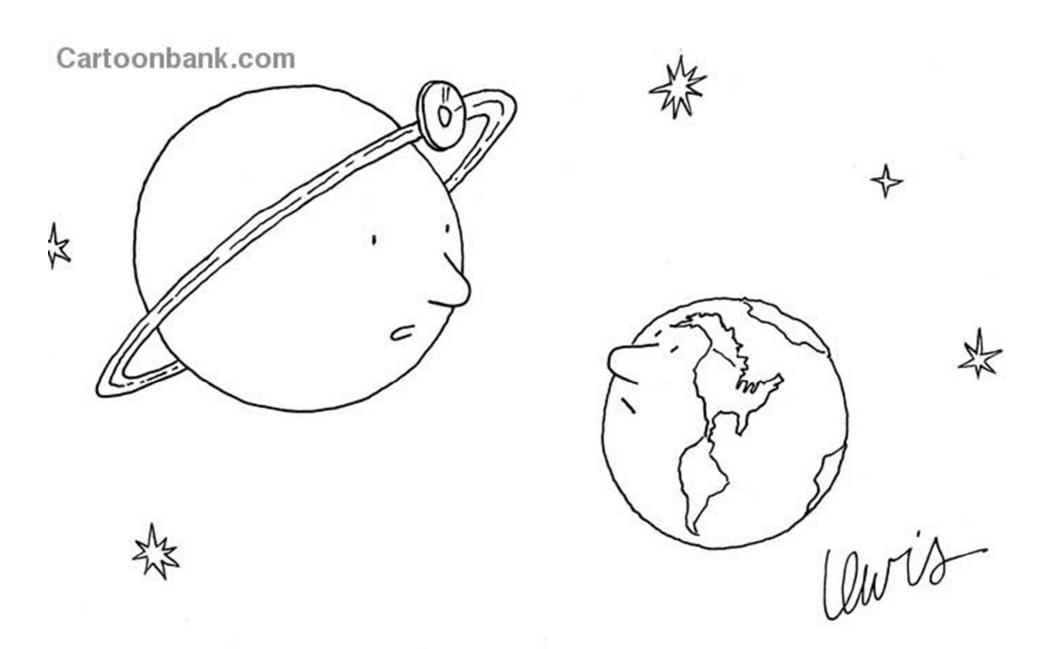


Conclusions

•The future will be interesting and surprising!

•We have work to do!





"Im afraid you have humans."